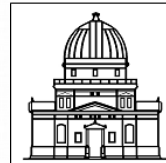


# Can I get a paper's worth of science done in one single Jupyter notebook?

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Dot Astronomy TO  
October 2019

KATHARINA LUTZ



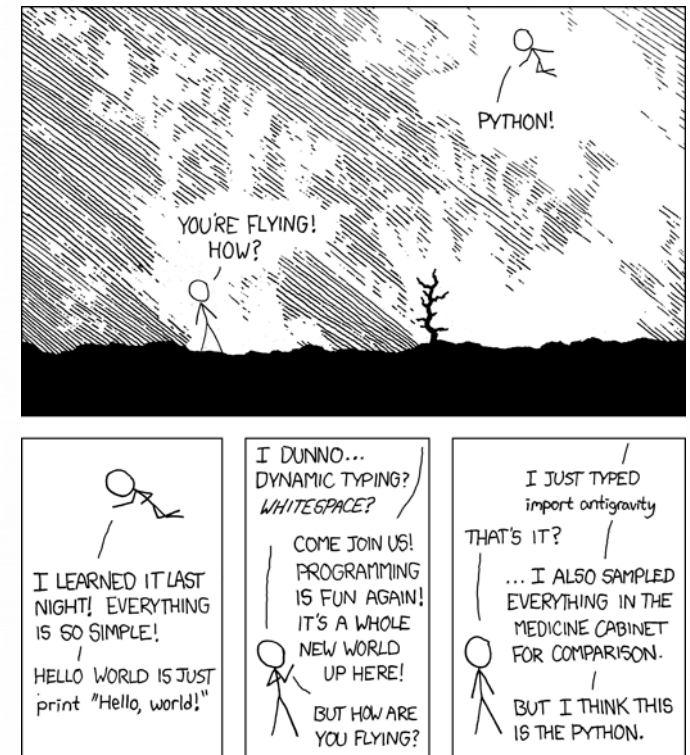
Observatoire **astronomique**  
de Strasbourg | ObAS





# Motivation

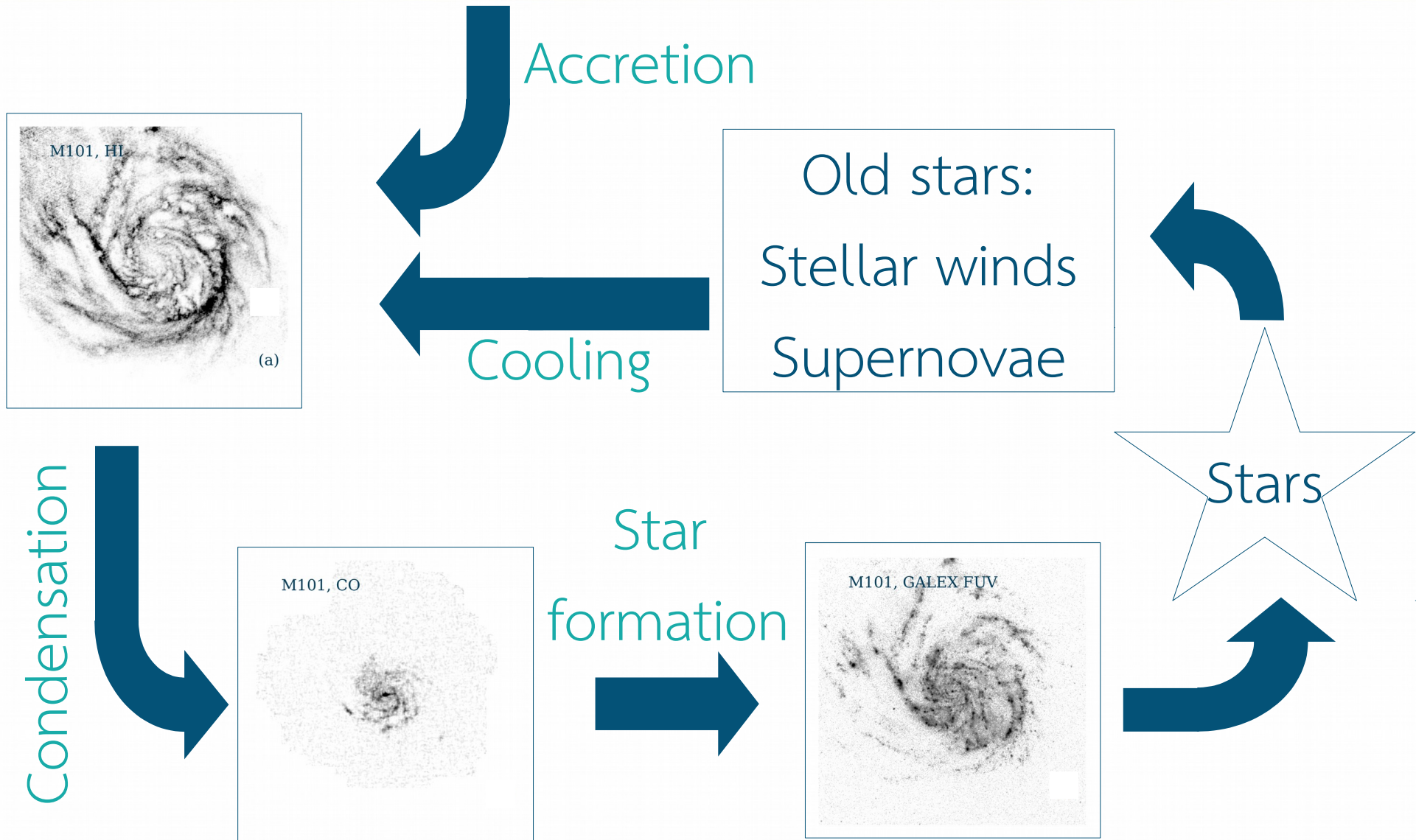
- Emergence of Science Platforms
- On offer: Jupyter hubs / notebooks
- Focus on Python because it's versatile
- Play with/ test CDS developments





# Short intro to the science

## Star formation – gas cycle



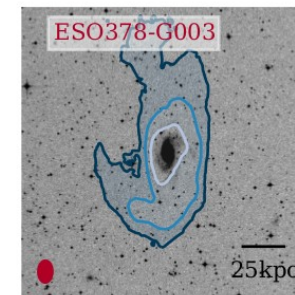
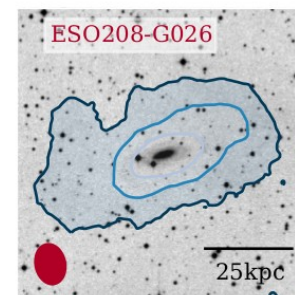
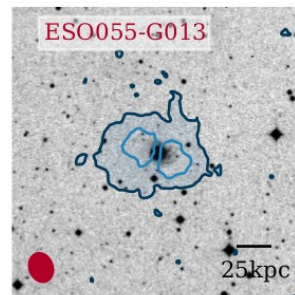
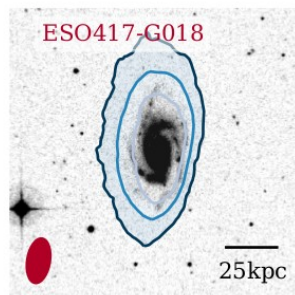


# Short intro to the science

## The HIX galaxy survey

Why do HI eXtreme galaxies have more massive HI discs than other galaxies?

- More gas accretion
- Less efficient star formation



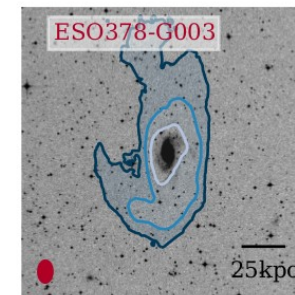
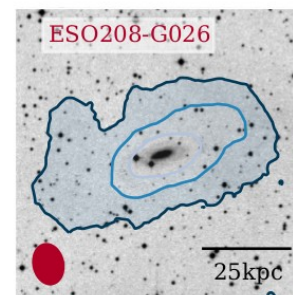
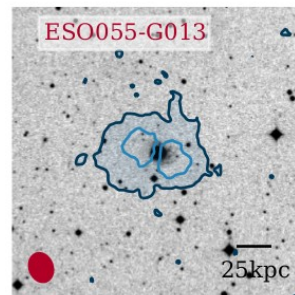
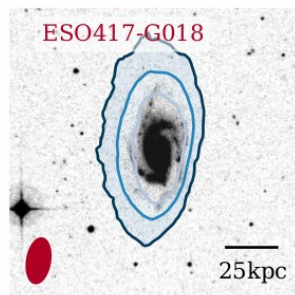


# Short intro to the science

## The HIX galaxy survey

Why do HI eXtreme galaxies have more massive HI discs than other galaxies?

- More gas accretion ???
- Less efficient star formation due to high spin / some gas does not contribute to star formation





# Quest

- We have observed optical images in g and r band with DECam for 4 HIX galaxies
- Now we want to:
  - Find optical CCD imaging for more galaxies
  - Measure photometry
  - Model the SEDs → dust and SF history
  - Compare results to more average galaxies



# Quest

Imagine I am working on a cloud space where I can access (VO) archives and do “pip install [any pypi package]”:

**Can I do the science for the next paper within one Jupyter notebook?**



# Getting Data

- ✓ No worries as long as data are available via astroquery or VO standards
- ✓ Lots of ongoing development (hips2fits)

# Getting Data

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J/AJ/157/81 : The Arecibo PPS Survey. I. Harvesting ALFALFA (O'Donoghue+, 2019)
J/AJ/157/194 : Total ALFALFA H I fluxes for extended sources (Hoffman+, 2019)
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2 dustpedia_set = tap_vizier.search("SELECT * FROM tables " +
3 "WHERE description LIKE '%Clark%' AND " +
4 "table_name LIKE '%A+A/609%'").to_table()
5 dustpedia_set['table_name', 'description']
```

Out[4]: Table masked=True length=4

table_name	description
J/A+A/609/A37/sample	Galaxy sample ( Clark C.J.R., Verstocken S., Bianchi S., Fritz J., Viaene S., Smith M.W.L., Baes M., Casasola V., Cassara L.P., Davies J.I., De Looze I., De Vis P., Evans R., Galametz M., Jones A.P., Lianou S., Madden S., Mosenkov A.V., Xilouris M.)
J/A+A/609/A37/apphot	Aperture-matched photometry for 27 bands (GALEX, SDSS, 2MASS, WISE, Spitzer, and Herschel) ( Clark C.J.R., Verstocken S., Bianchi S., Fritz J., Viaene S., Smith M.W.L., Baes M., Casasola V., Cassara L.P., Davies J.I., De Looze I., De Vis P., Evans R., Galametz M., Jones A.P., Lianou S., Madden S., Mosenkov A.V., Xilouris M.)
J/A+A/609/A37/irasphot	Ancillary photometry from IRAS SCANPI ( Clark C.J.R., Verstocken S., Bianchi S., Fritz J., Viaene S., Smith M.W.L., Baes M., Casasola V., Cassara L.P., Davies J.I., De Looze I., De Vis P., Evans R., Galametz M., Jones A.P., Lianou S., Madden S., Mosenkov A.V., Xilouris M.)
J/A+A/609/A37/plnphot	Ancillary photometry from Planck CCS2 ( Clark C.J.R., Verstocken S., Bianchi S., Fritz J., Viaene S., Smith M.W.L., Baes M., Casasola V., Cassara L.P., Davies J.I., De Looze I., De Vis P., Evans R., Galametz M., Jones A.P., Lianou S., Madden S., Mosenkov A.V., Xilouris M.)

```
In [5]: 1 dustpedia = tap_vizier.search("SELECT * FROM \"J/A+A/609/A37/sample\").to_table()
2 dustpedia_photo = tap_vizier.search("SELECT * " +
3 "FROM \"J/A+A/609/A37/apphot\").to_table()
4
5 for col in dustpedia_photo.colnames:
6     if col not in dustpedia.colnames:
7         dustpedia[col] = dustpedia_photo[col]
8 dustpedia.write('Tables/dustpedia.vot', format='votable', overwrite=True)
9 dustpedia = Table.read('Tables/dustpedia.vot')
10 dustpedia
```

Out[5]: Table masked=True length=875

recno	Name	RAJ2000	DEJ2000	MType	MClass	D25	Incl	HV	HVcorr	Distcorr	Dist0	HVH	DistH	Dist0H	Dist	Simb
		deg	deg			arcmin	deg	km / s	km / s	Mpc	Mpc	km / s	Mpc	Mpc	Mpc	
int32	object	float64	float64	float64	bytes4	float64	float64	int16	int32	float64	float64	int32	float64	float64	float64	
41	NGC0678	27.35355	21.99728	3.0	SBb	3.0903	90.0	2835	2766	37.76625	27.87	2793	38.1349	--	27.87	N
44	NGC0855	33.51495	27.87685	-4.8	E	2.9512	90.0	592	693	9.46204	9.28	592	8.08301	9.63829	9.63829	N
52	NGC1056	40.70145	28.57372	1.0	Sa	1.8621	46.700001	1545	1561	21.31349	26.8	1544	21.08138	31.18893	31.18893	N
56	UGC02392	43.9434	33.76658	5.9	Sc	1.3183	76.599998	1548	1583	21.61387	26.52	1546	21.10868	--	26.52	UC
125	UGC03016	65.33325	36.76014	5.9	Sc	1.122	90.0	5897	5806	79.27362	30.05	2458	33.5609	--	30.05	UC
48	NGC0925	36.82035	33.579	7.0	Scd	10.7152	58.700001	553	676	9.22993	8.61	554	7.56417	9.20449	9.20449	N



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J/A+A/609/A37/pinkphot	Ancillary photometry from Planck CCS2 (Clark C.J.R., Verstocken S., Bianchi S., Fritz J., Viaene S., Smith M.W.L., Baes M., Casasola V., Cassara L.P., Davies J.I., De Looze I., De Vis P., Evans R., Galametz M., Jones A.P., Lianou S., Madden S., Mosenkov A.V., Xilouris M.)

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```
In [6]: 1 dp_alf = XMatch.query(cat1=open('Tables/dustpedia.vot'),
        2                       cat2='vizier:J/ApJ/861/49/table2',
        3                       max_distance=1.75 * u.arcmin, colRA1='RAJ2000',
        4                       colDecl1='DEJ2000')
        5 dp_alf
```

Out[6]: Table masked=True length=343

angDist	recno_1	Name_1	RAJ2000_1	DEJ2000_1	MType	MClass	D25	Incl	HV	HVcorr	Distcorr	Dist0	HVH	DistH	Dist0H
float64	int64	str23	float64	float64	float64	str4	float64	float64	int64	int64	float64	float64	int64	float64	float64
14.287441	41	NGC0678	27.35355	21.99728	3.0	SBb	3.0903	90.0	2835	2766	37.76625	27.87	2793	38.1349	--
31.099272	44	NGC0855	33.51495	27.87685	-4.8	E	2.9512	90.0	592	693	9.46204	9.28	592	8.08301	9.63829
8.557775	52	NGC1056	40.70145	28.57372	1.0	Sa	1.8621	46.700001	1545	1561	21.31349	26.8	1544	21.08138	31.18893



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        3                                     "WHERE description LIKE '%Clarks' AND " +
        4                                     "table name LIKE '%A+/609%'").to_table()
        5 dustpedia_set['table_name', 'description']
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table_name	description
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J/A+A/609/A37/apphot	Aperture-matched photometry for 27 bands (GALEX, SDSS, 2MASS, WISE, Spitzer, and Herschel) (Clark C.J.R., Verstocken S., Bianchi S., Fritz J., Viaene S., Smith M.W.L., Baes M., Casasola V., Cassara L.P., Davies J.I., De Looze I., De Vis P., Evans R., Galametz M., Jones A.P., Lianou S., Madden S., Mosenkov A.V., Xilouris M.)
J/A+A/609/A37/irasphot	Ancillary photometry from IRAS SCANPI (Clark C.J.R., Verstocken S., Bianchi S., Fritz J., Viaene S., Smith M.W.L., Baes M., Casasola V., Cassara L.P., Davies J.I., De Looze I., De Vis P., Evans R., Galametz M., Jones A.P., Lianou S., Madden S., Mosenkov A.V., Xilouris M.)
J/A+A/609/A37/pinkphot	Ancillary photometry from Planck CCS2 (Clark C.J.R., Verstocken S., Bianchi S., Fritz J., Viaene S., Smith M.W.L., Baes M., Casasola V., Cassara L.P., Davies J.I., De Looze I., De Vis P., Evans R., Galametz M., Jones A.P., Lianou S., Madden S., Mosenkov A.V., Xilouris M.)

```
In [5]: 1 dustpedia = tap_vizier.search("SELECT * FROM \"J/A+/609/A37/sample\"").to_table()
        2 dustpedia_photo = tap_vizier.search("SELECT * " +
        3                                     "FROM \"J/A+/609/A37/apphot\"").to_table()
        4
        5 for col in dustpedia_photo.colnames:
        6     if col not in dustpedia.colnames:
        7         dustpedia[col] = dustpedia_photo[col]
        8 dustpedia.write('Tables/dustpedia.vot', format='votable', overwrite=True)
        9 dustpedia = Table.read('Tables/dustpedia.vot')
        10 dustpedia
```

Out[5]: Table masked=True length=875

recno	Name	RAJ2000	DEJ2000	MType	MClass	D25	Incl	HV	HVcorr	Distcorr	Dist0	HVH	DistH	Dist0H	Dist	Slmb
		deg	deg			arcmin	deg	km / s	km / s	Mpc	Mpc	km / s	Mpc	Mpc	Mpc	
int32	object	float64	float64	float64	bytes4	float64	float64	int16	int32	float64	float64	int32	float64	float64	float64	
41	NGC0678	27.35355	21.99728	3.0	SBb	3.0903	90.0	2835	2766	37.76625	27.87	2793	38.1349	--	27.87	N
44	NGC0855	33.51495	27.87685	-4.8	E	2.9512	90.0	592	693	9.46204	9.28	592	8.08301	9.63829	9.63829	N

```
In [6]: 1 dp_alf = XMatch.query(cat1=open('Tables/dustpedia.vot'),
        2                       cat2='vizier:J/ApJ/861/49/table2',
        3                       max_distance=1.75 * u.arcmin, colRA1='RAJ2000',
        4                       colDec1='DEJ2000')
        5 dp_alf
```

Out[6]: Table masked=True length=343

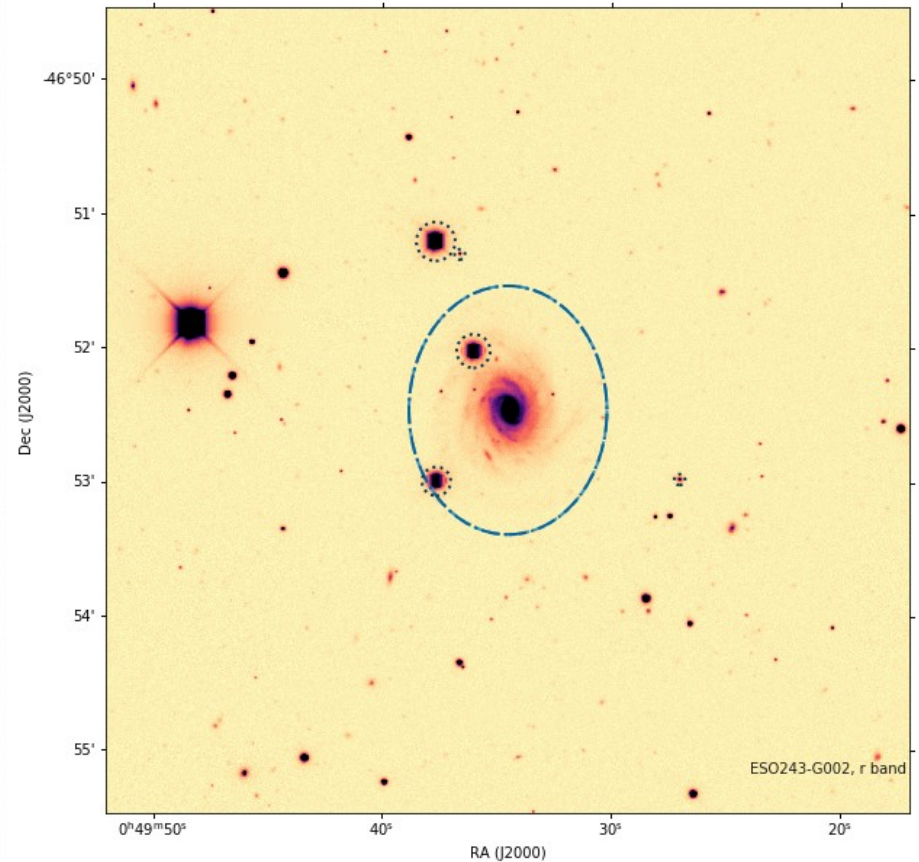
angDist	recno_1	Name_1	RAJ2000_1	DEJ2000_1	MType	MClass	D25	Incl	HV	HVcorr	Distcorr	Dist0	HVH	DistH	Dist0H
float64	int64	str23	float64	float64	float64	str4	float64	float64	int64	int64	float64	float64	int64	float64	float64
14.287441	41	NGC0678	27.35355	21.99728	3.0	SBb	3.0903	90.0	2835	2766	37.76625	27.87	2793	38.1349	--
31.099272	44	NGC0855	33.51495	27.87685	-4.8	E	2.9512	90.0	592	693	9.46204	9.28	592	8.08301	9.63829
8.557775	52	NGC1056	40.70145	28.57372	1.0	Sa	1.8621	46.70001	1545	1561	21.31349	26.8	1544	21.08138	31.18893



# Measuring photometry

## ✓ Photutils works

- Find stars
- Measure aperture photometry
- Mask stars
- Visualise aperture for double-checking



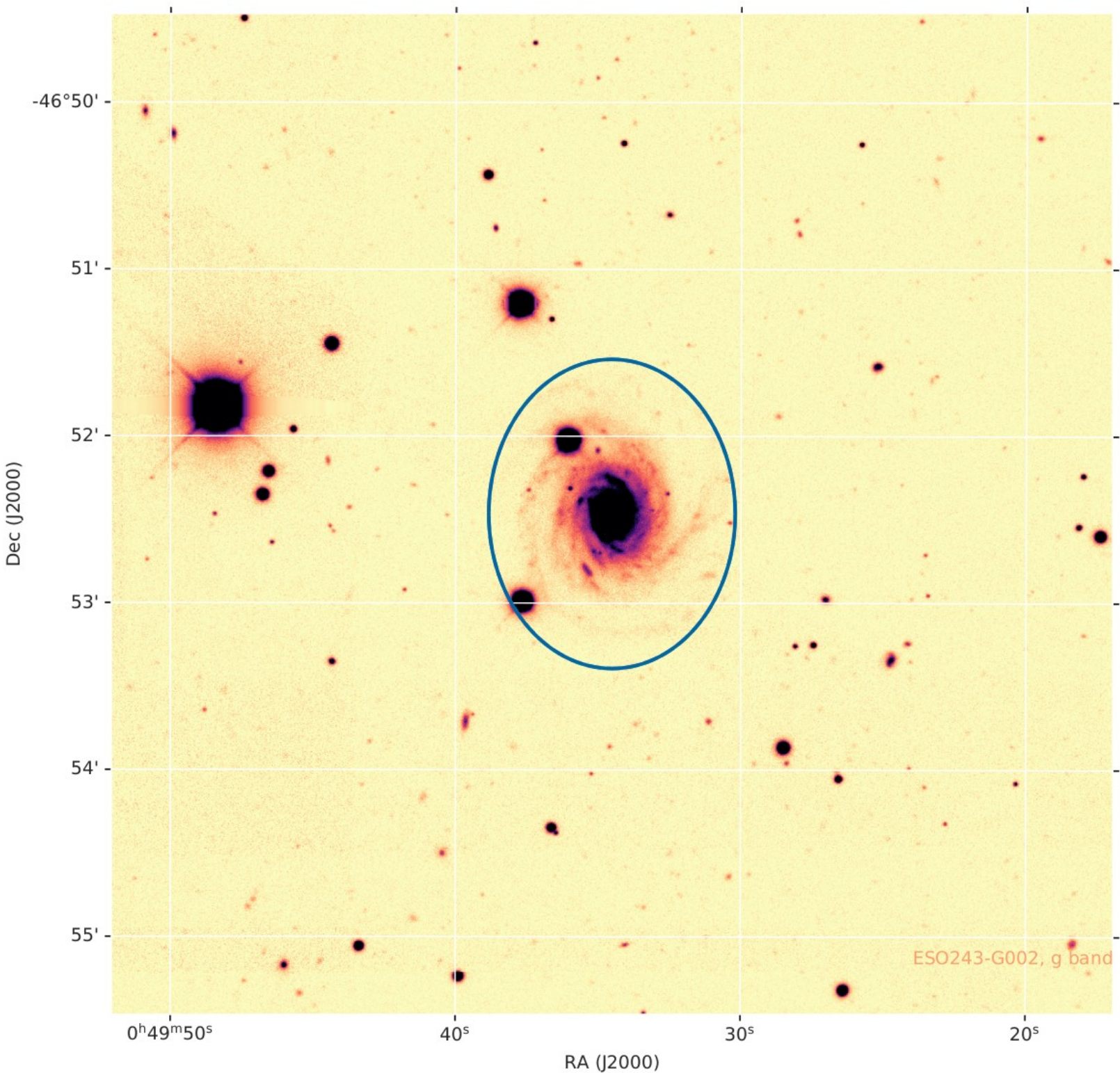


# Visualisation

## ✓ Matplotlib PLUS

- Plot functions for special objects
- WCS axes
- `astropy.visualisation`

## ✓ Aladin Lite widget



23/10/



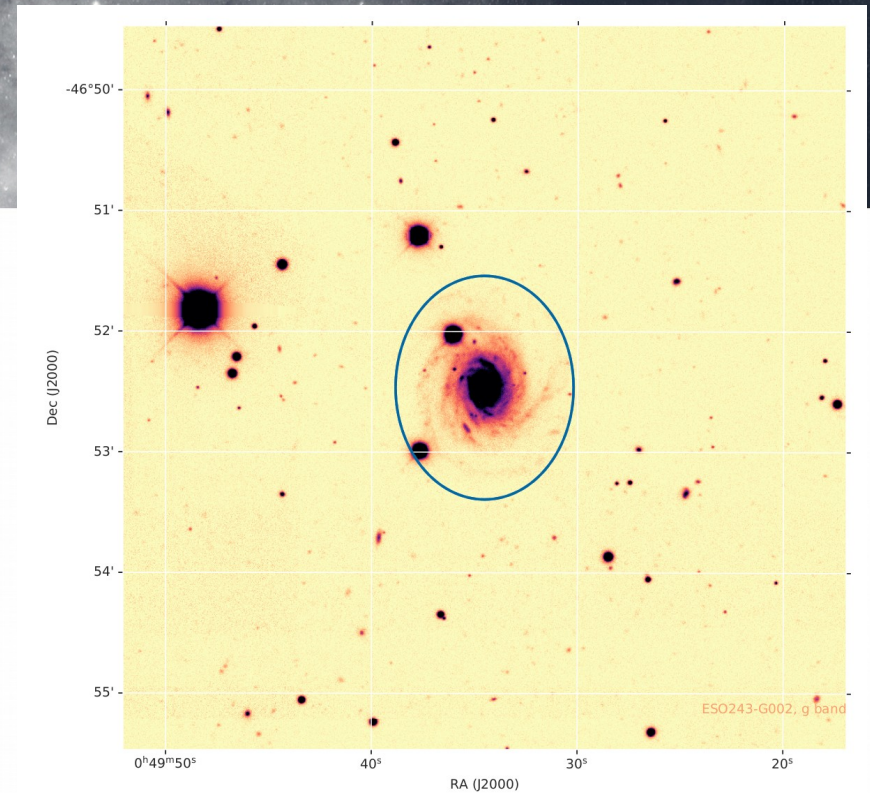


# Visualisation

## ✓ Matplotlib PLUS

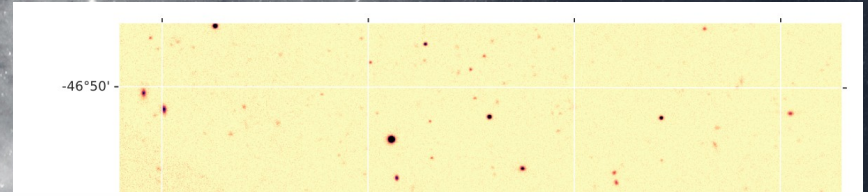
- Plot functions for special objects
- WCS axes
- astropy.visualisation

## ✓ Aladin Lite widget





# Visualisation



```
In [7]: 1 aladin1 = ipyal.Aladin(survey='P/DSS2/color', fov=0.1,  
2                               target='ES0243-G002')  
3 aladin1
```

J2000 00 49 34.522 -46 52 27.94

FoV: 5.99'

ALADIN

ALADIN: THE VISUAL

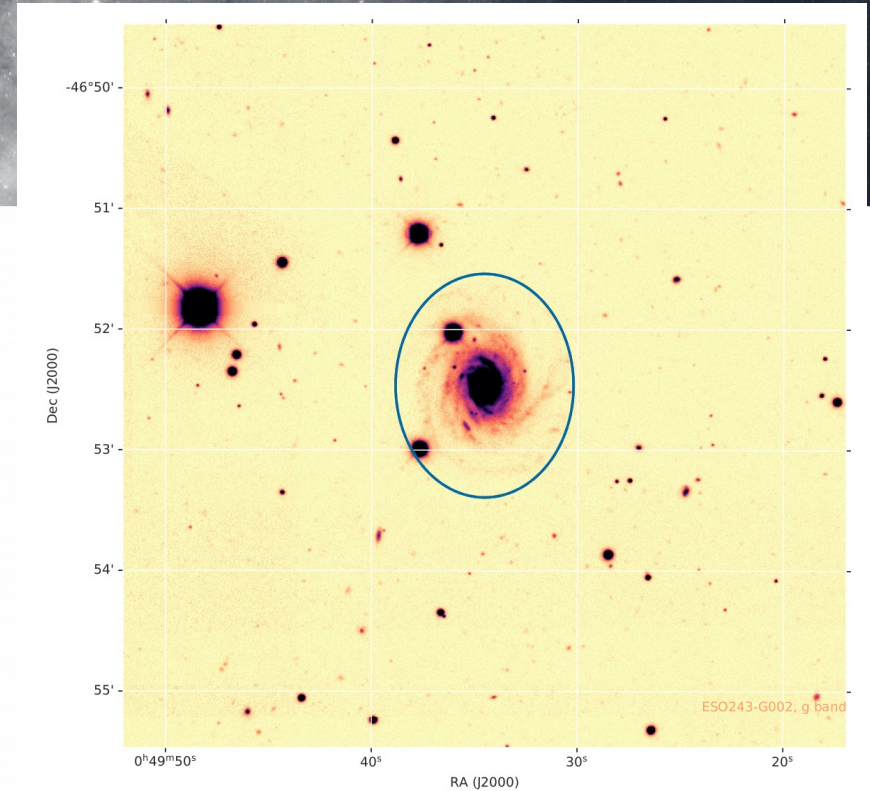


# Visualisation

## ✓ Matplotlib PLUS

- Plot functions for special objects
- WCS axes
- astropy.visualisation

## ✓ Aladin Lite widget

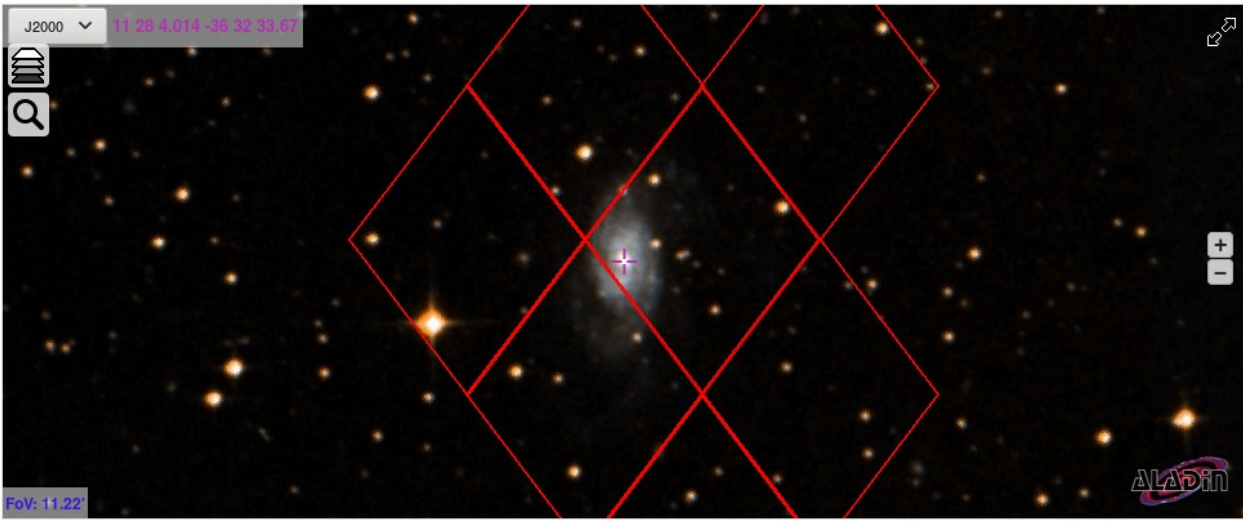




# CDS developments in context

- Aladin Lite widget nice for quick visualisation
- MOC help to avoid source confusion

```
In [6]: 1 aladin= ipyal.Aladin(survey='P/DSS2/color', target='ES0378-3', fov=0.5)
        2 aladin
```

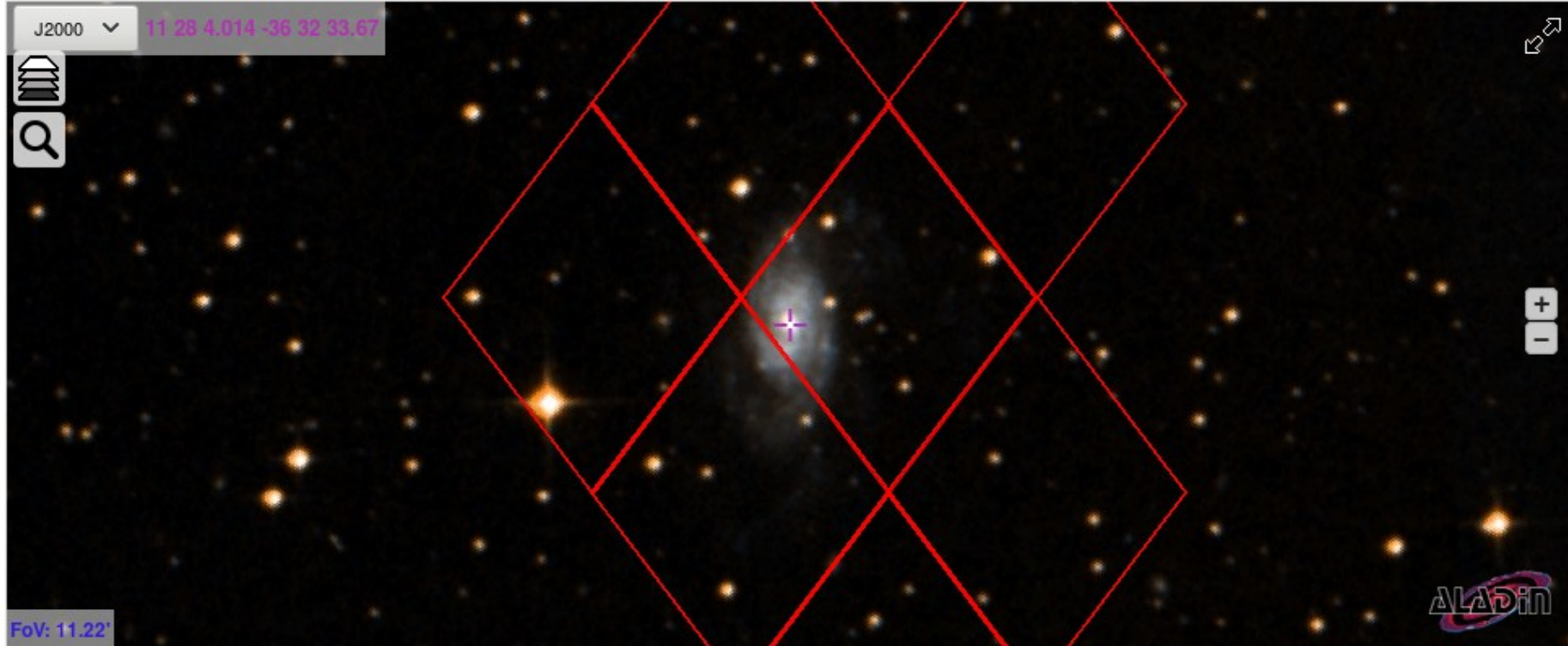


```
In [9]: 1 aladin.add_moc_from_dict(my_moc.serialize(format='json'), {'color': 'red'})
```

# CDS developments in context

## Aladin Lite widget nice for quick visualization

```
In [6]: 1 aladin= ipyal.Aladin(survey='P/DSS2/color', target='ES0378-3', fov=0.5)
        2 aladin
```



```
In [9]: 1 aladin.add_moc_from_dict(my_moc.serialize(format='json'), {'color': 'red'})
```



# SED modelling

- ⊗ Both MAGPHYS and CIGALE are command line tools
- ⊙ Maybe the astropy modelling + table of model?



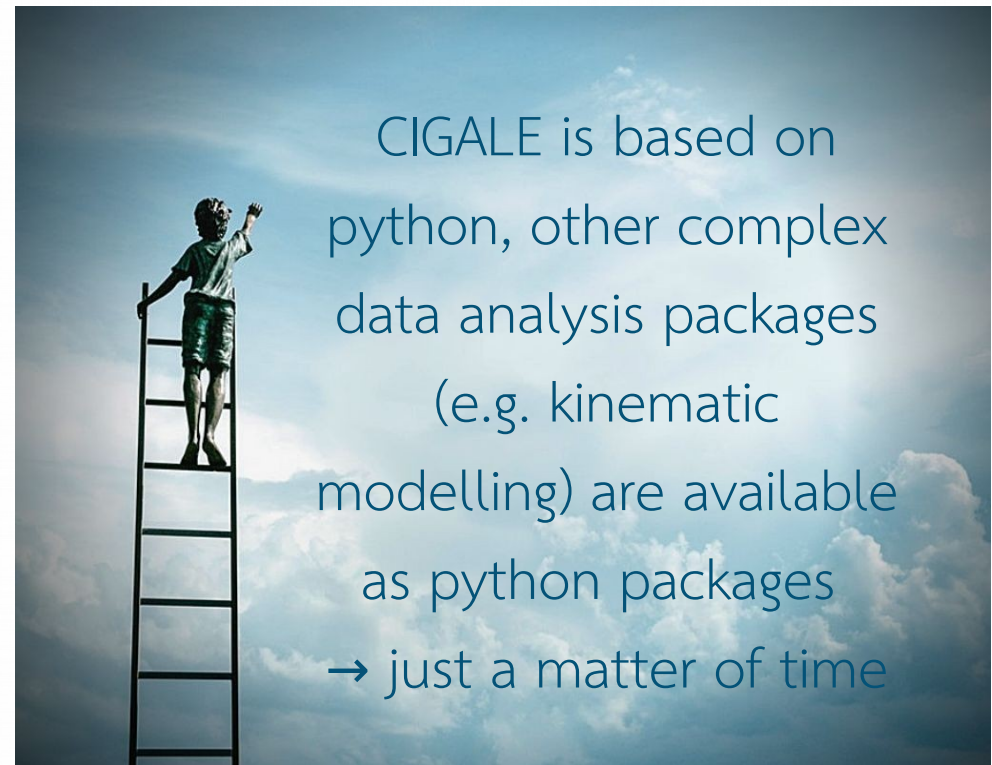
# SED modelling

- ⊗ Both MAGPHYS and CIGALE are command line tools
- ⊙ Maybe the astropy modelling + table of model? Happy to hear more ...



# SED modelling

- ⊗ Both MAGPHYS and CIGALE are command line tools
- ? Maybe the astropy modelling + table of model? Happy to hear more...





# Quest

Imagine I am working on a cloud space where I can access the VO resources and do “pip install [any pypi package]”:

Can I do the science for the next paper within one Jupyter notebook?

**MOSTLY**



# Conclusion & Outlook

- Lots of heavy lifting can already be done in Jupyter notebooks:
  - Background measurements
  - Data cube analysis
- Many good things are on the best way:
  - Hips2fits module for astroquery (in development)
  - ESAsky widget for Jupyter notebooks (pyesasky)
- Crucial: access to outside databases!